



Workflows e Plataformas Científicas: preparação para o LSST DR1

25 e 26 de março de 2025 | das 13h às 17h



Projeto de uma plataforma-ISSC-LSST para classificação automática de galáxias com aprendizagem profunda

Reinaldo R. Rosa

Coordenadoria de Pesquisa Aplicada e Desenvolvimento Tecnológico (COPDT)
INPE-MCTI

LSST-Sci. Colab.

1. Galaxies SC
2. S-MW-LV SC
3. Solar System SC
4. Dark Energy SC
5. AGN SC
6. Transients & VS SC
7. Strong L. SC
8. Inf. & Stat. SC



Thales



Luan



Vitor



Reinaldo



Rubens



Carlos E. Falandes

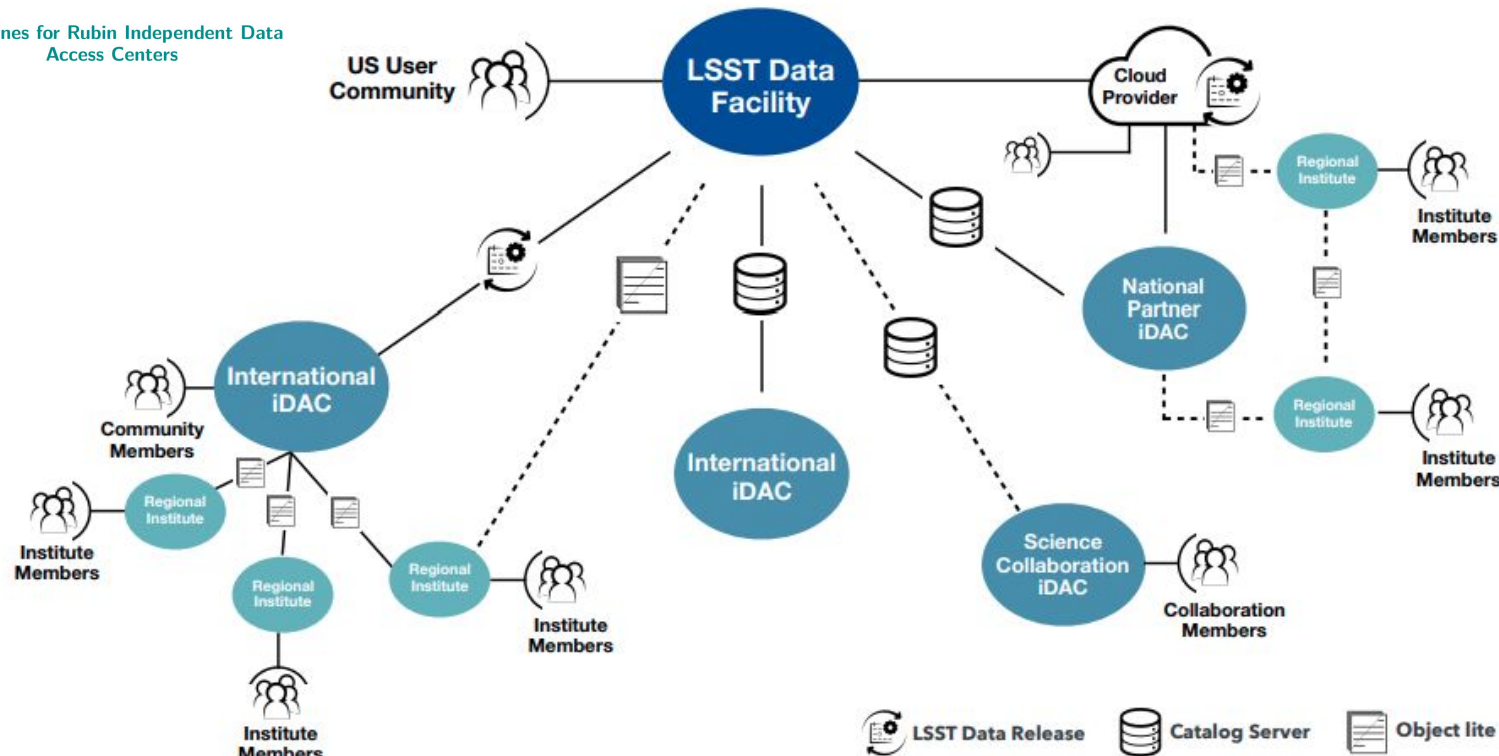
Data Science Fellowship Program



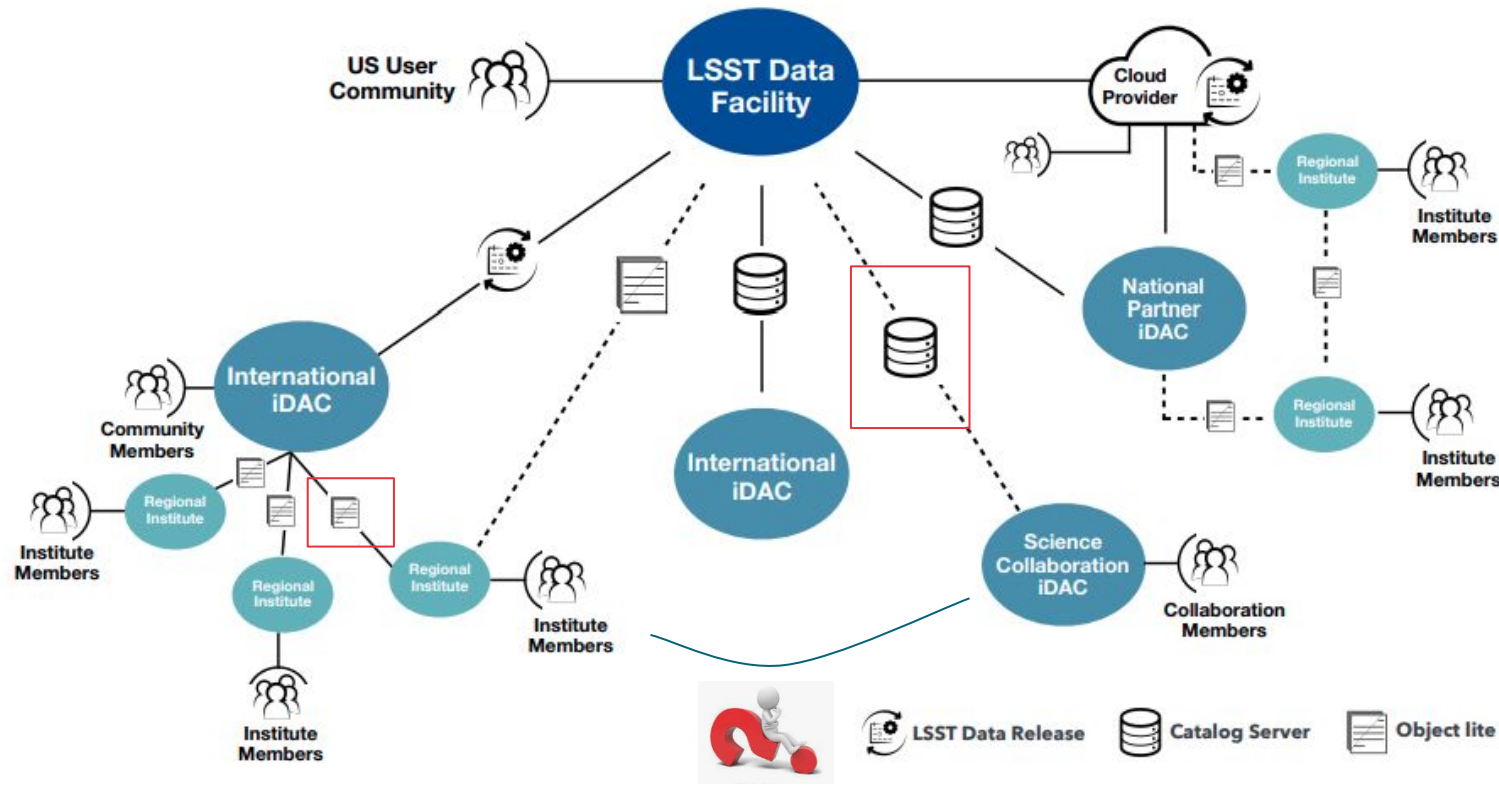
LSST
Discovery Alliance

IDAC network topology

Guidelines for Rubin Independent Data
Access Centers



IDAC network topology



ISSC liaisons:

- [LSST Survey Cadence Optimization Committee](#): Federica Bianco (University of Delaware)
- Rubin Data Management: Leanne Guy (Rubin)



ISSC Executive Council:

ISSC co-Chairs:

- Francois Lanusse (CNRS)
- Ashley Villar (Harvard)

- Melissa DeLucchi
- Alex Gagliano
- Matthew Graham
- Francois Lanusse (ex-officio)
- Alex Malz
- Ashish Mahabal
- Ashley Villar (ex-officio)

ISSC Leadership Team

Ethics Committee:

- Eric Aubourg
- Reinaldo Roberto Rosa
- Lior Shamir (chair)
- Yuanyuan Zhang

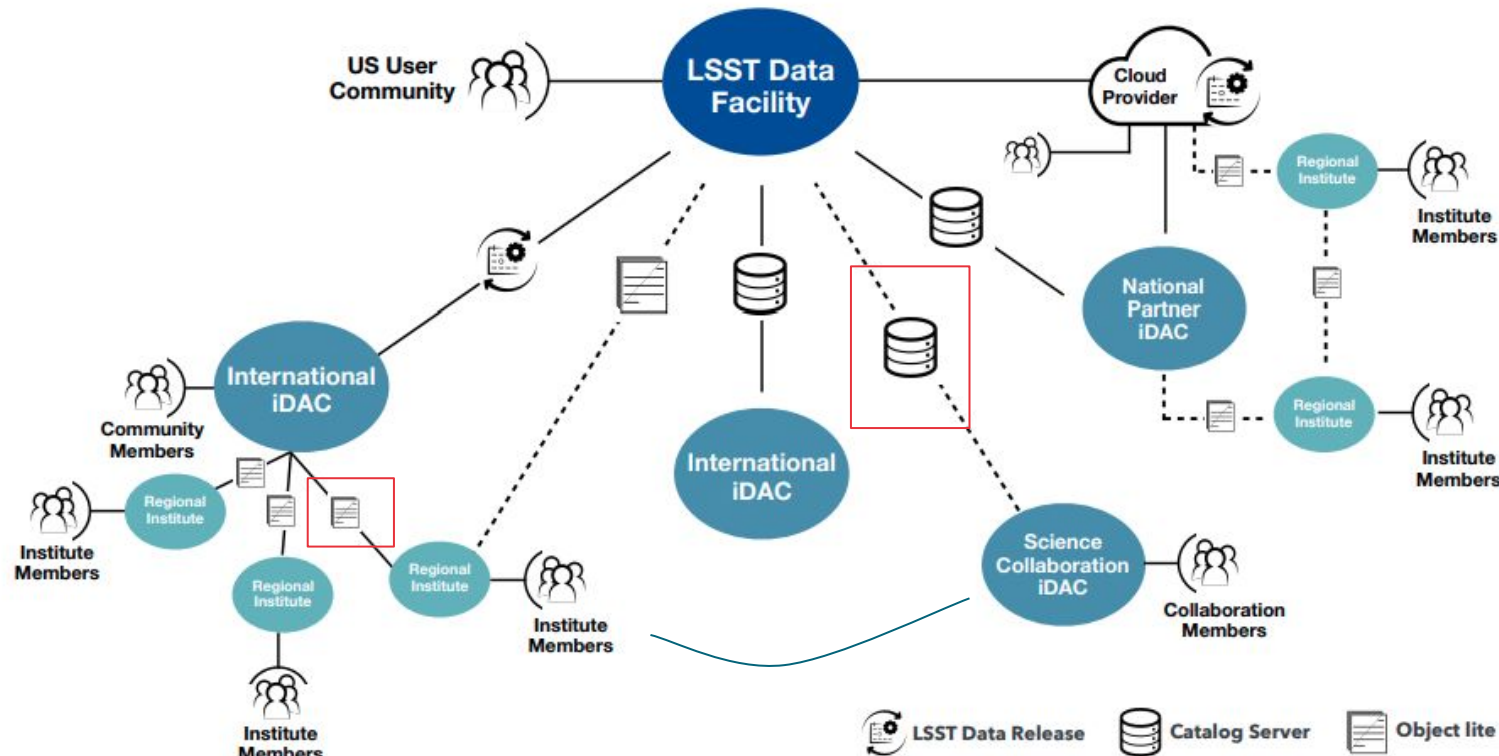
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- Alex Malz
- Josh Speagle

Membership Committee:

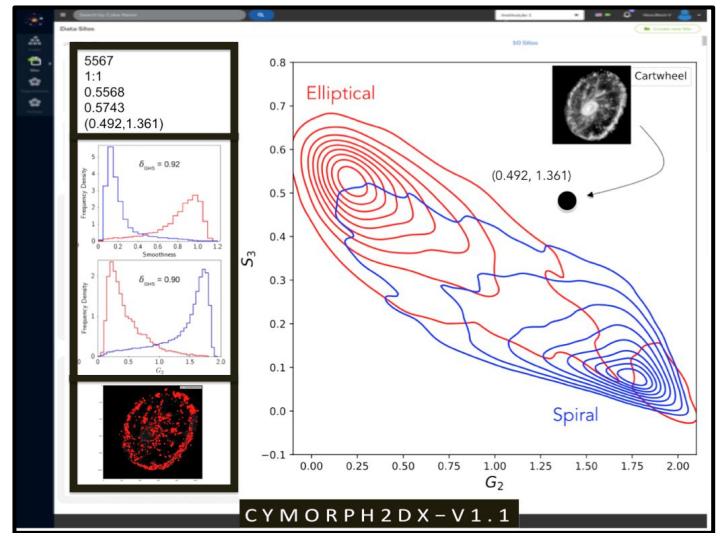
- Tatiana Acero-Cuellar
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- Bryce Kalmbach (chair)
- Konstantin Malanchev

IDAC network topology



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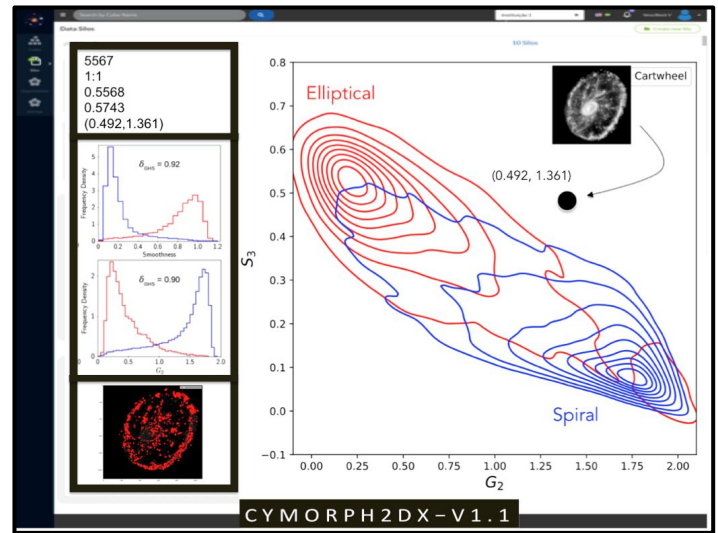
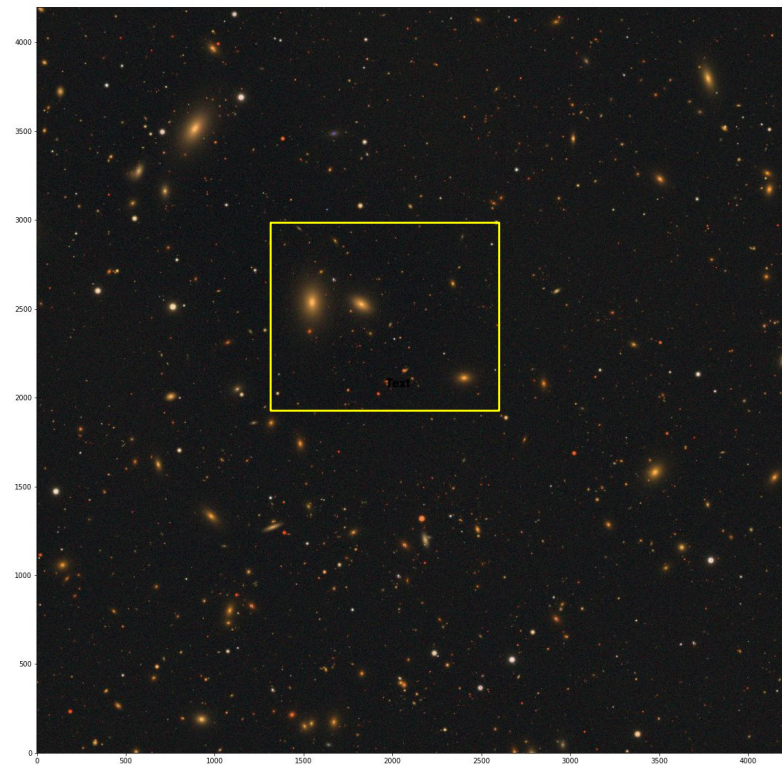
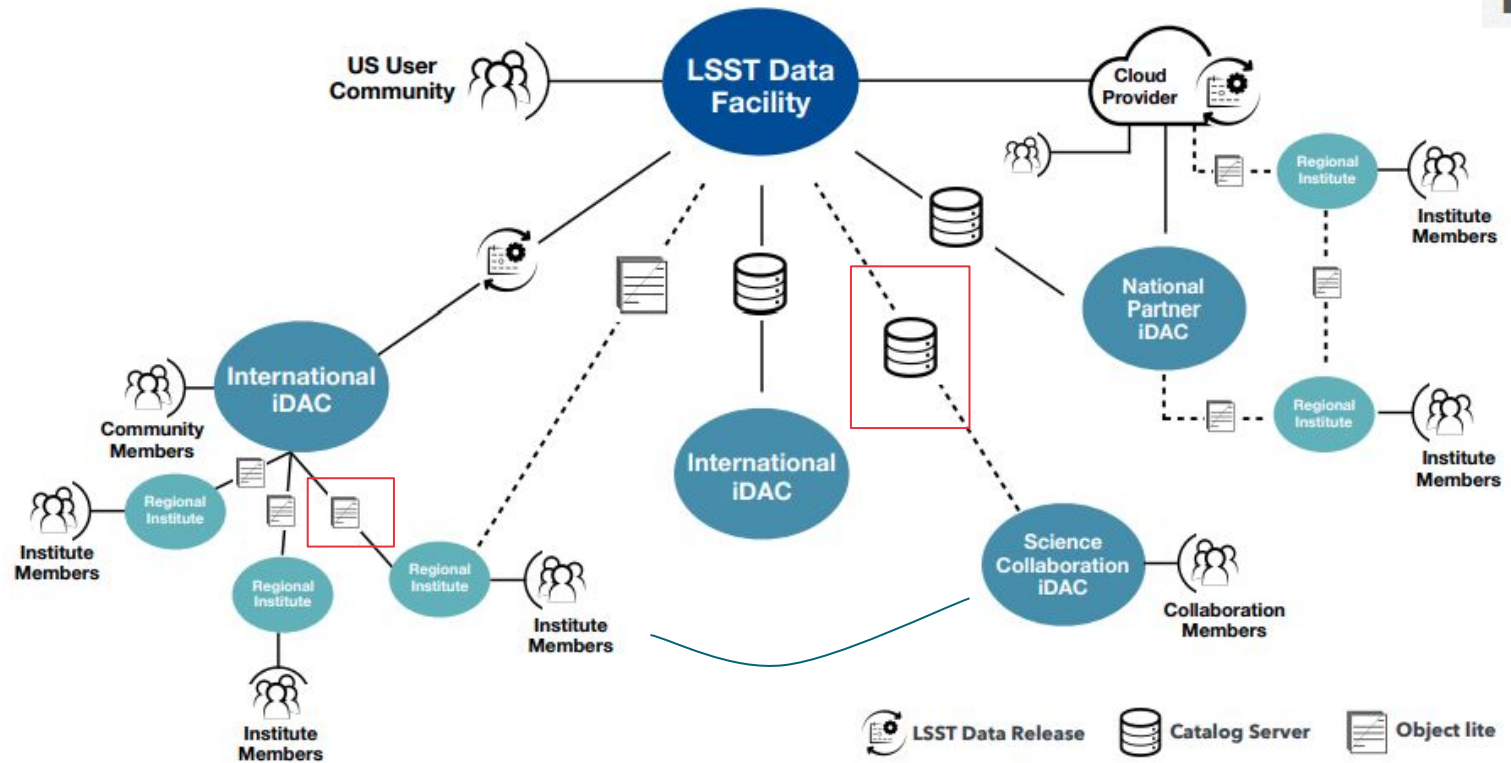
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Vera C. Rubin Observatory Legacy Survey of Space and Time

Informatics & Statistics Science Collaboration

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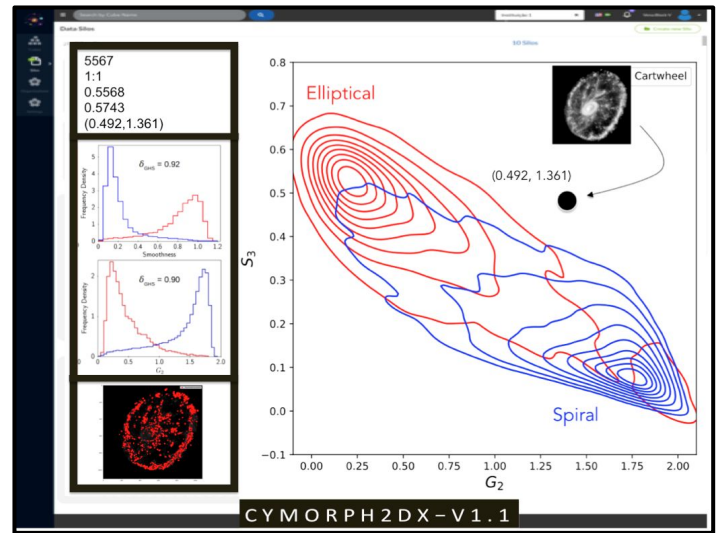
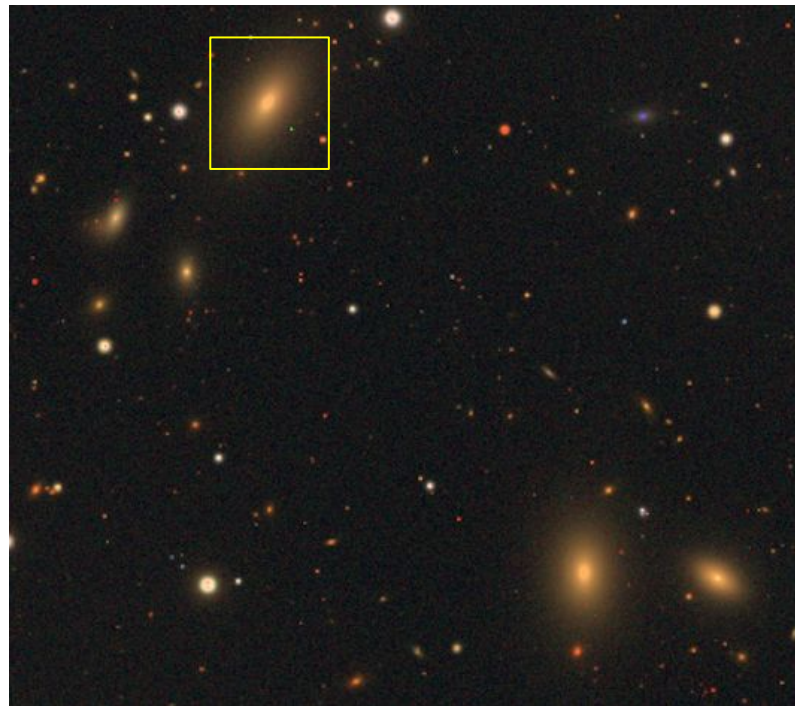
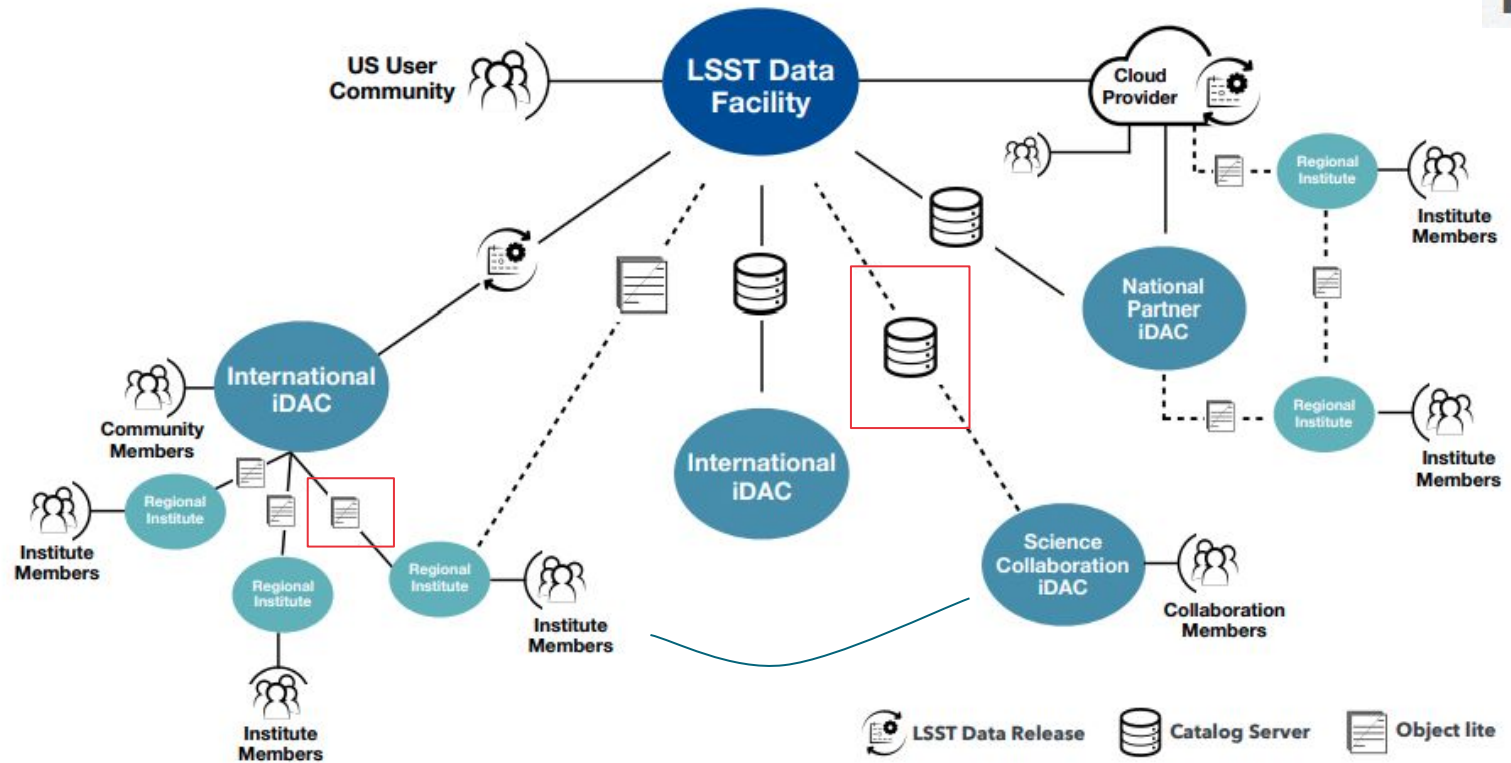
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User Generated DP 100% baseado na nuvem?

RH/\$: ISSC e/ou IDAC-BPG?

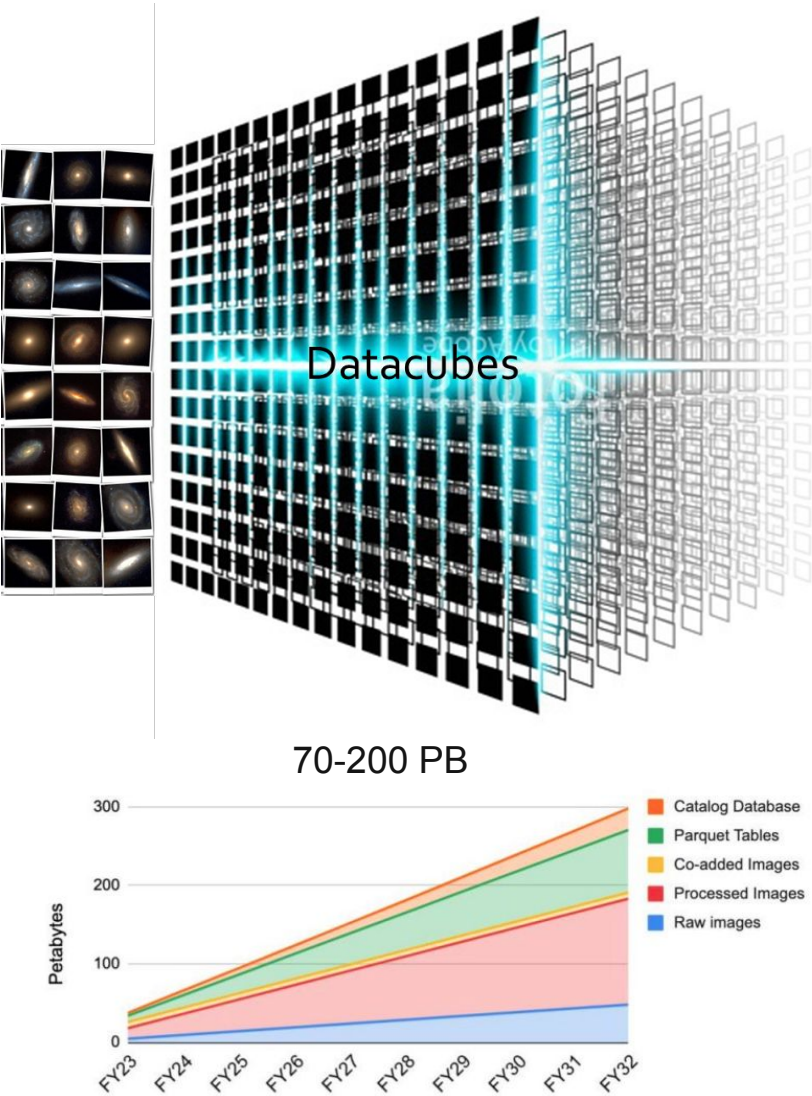
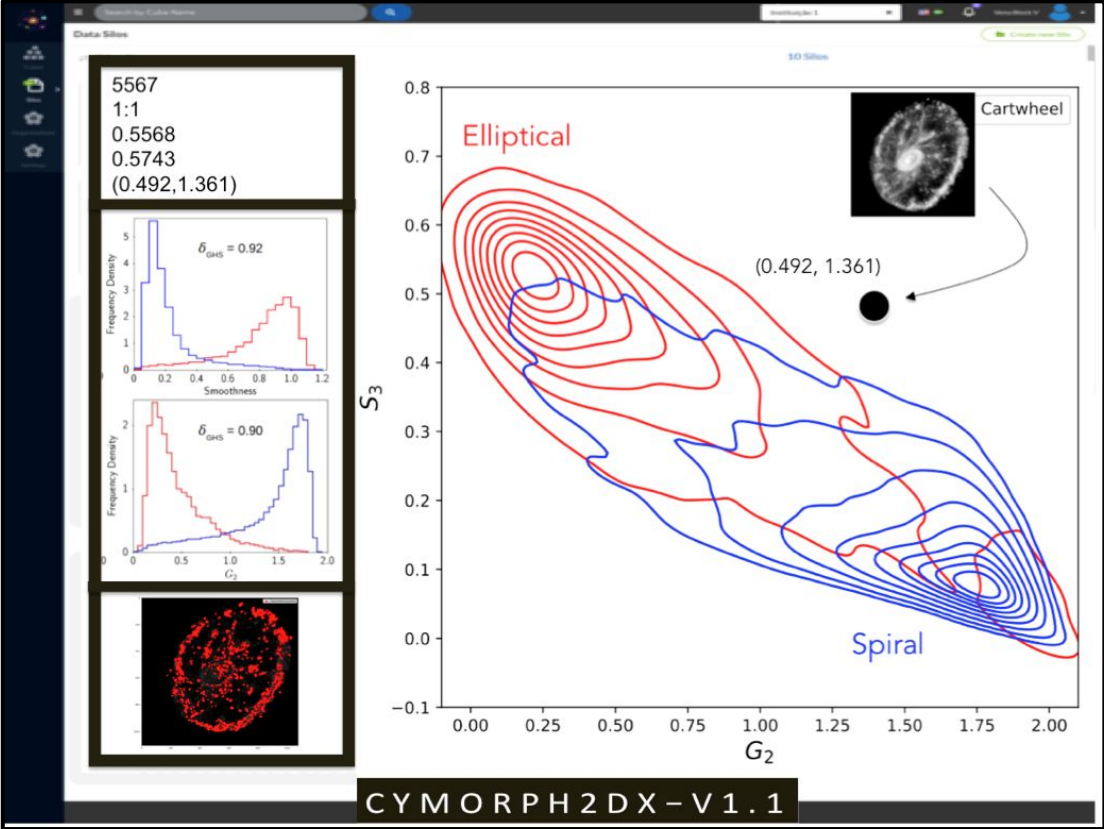


FIGURE 1: Volume of different datatypes over time from DMTN-135 data.

Guidelines for Rubin Independent Data
Access Centers

Morphometrics

CAS

$$C = 5 \times \log \left(\frac{R_{80}}{R_{20}} \right)$$



$$A(\phi) = \frac{\sum |I_{\text{original}} - I_{\text{rotated}}|}{\sum I_{\text{original}}} - A_{\text{bkg}}$$



$$S = 10 \times \left(\frac{\sum |I_{\text{original}} - I_{\text{smoothed}}|}{\sum I_{\text{original}}} \right) - S_{\text{bkg}}$$



EGG

$$E(I) = - \sum_k^K p(I_k) \log[p(I_k)],$$



$$\tilde{E}(I) = \frac{E(I)}{E_{\text{max}}}, \text{ where } 0 \leq \tilde{E}(I) \leq 1 \text{ and } E_{\text{max}} = \log(K)$$

$$G = \frac{1}{2\bar{X}n(n-1)} \sum_{i=1}^n \sum_{j=1}^n |X_i - X_j|,$$



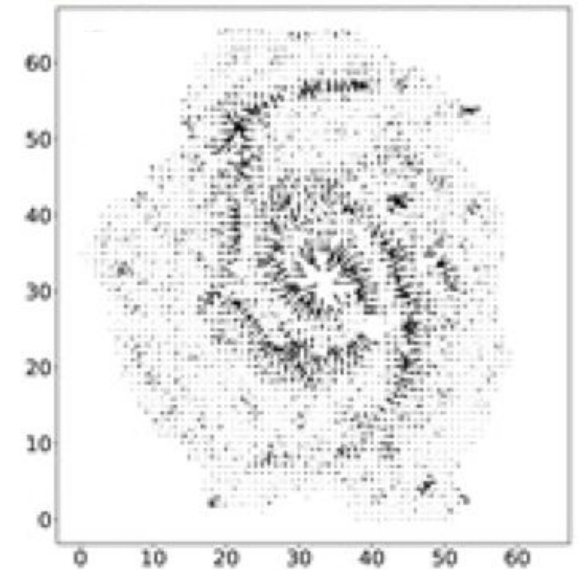
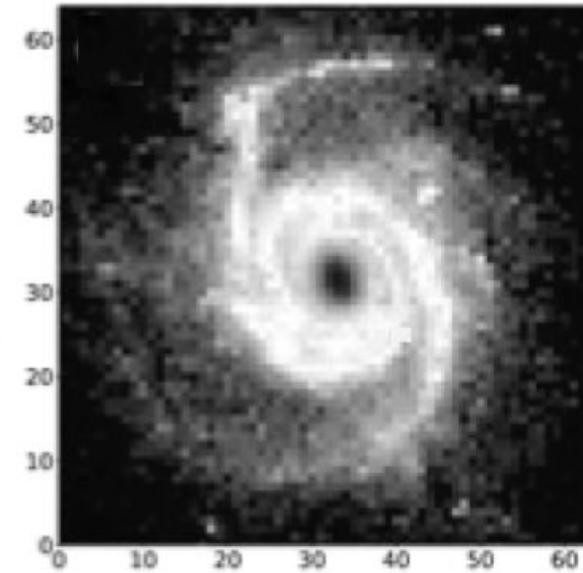
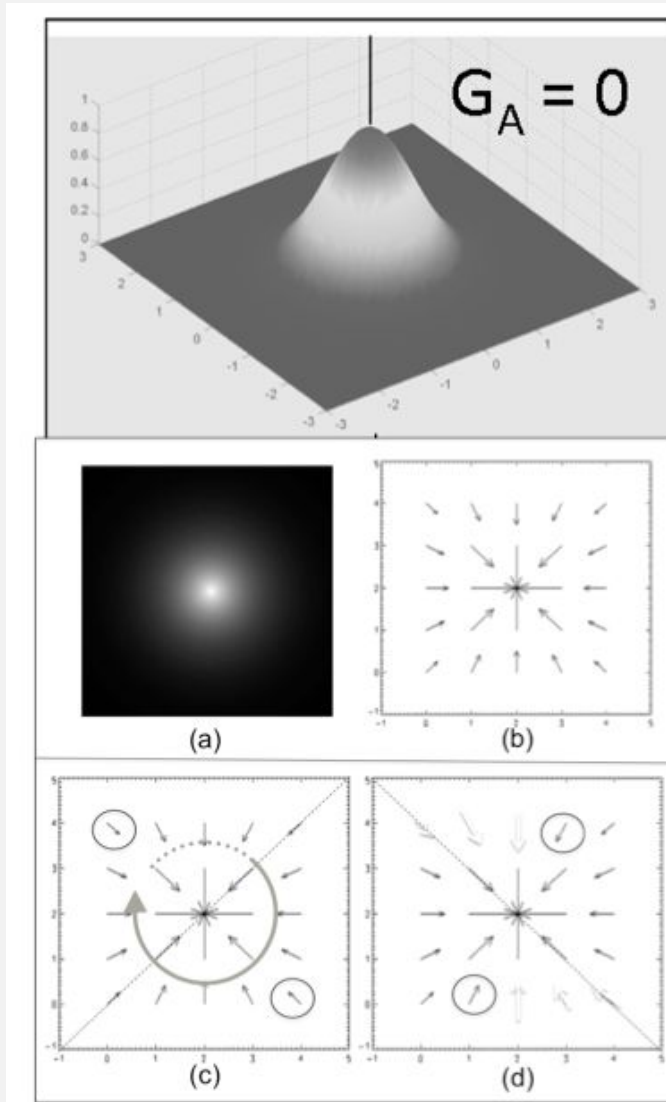
(total homogeneity) $0 < G < 1$ (total inh.)

$$G_2 = \frac{V_A}{V} (1 - \text{confluence})$$

$$\text{confluence} = \left(\frac{|\sum_i^{V_A} v_a^i|}{\sum_i^{V_A} |v_a^i|} \right)$$



+ G_1, G_3, G_4



$$G_2 = \frac{V_A}{N} \left(2 - \frac{|\sum_i^{V_A} r_i|}{\sum_i^{V_A} |r_i|} \right)$$

<https://doi.org/10.1093/mnras/sly054>

Monthly Notices
of the
ROYAL ASTRONOMICAL SOCIETY
MNRAS **477**, L101–L105 (2018)
Advance Access publication 2018 April 04
doi:10.1093/mnras/sly054

Gradient pattern analysis applied to galaxy morphology

R. R. Rosa,¹ R. R. de Carvalho,² R. A. Sautter,¹ P. H. Barchi,¹ D. H. Stalder,^{1,3} T. C. Moura,² S. B. Rembold,⁴ D. R. F. Morell⁵ and N. C. Ferreira¹

¹Lab for Computing and Applied Mathematics, National Institute for Space Research (INPE), S. J. dos Campos, 12245-970, Brazil

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³NIDTEC/FP-UNA, San Lorenzo, 2169-000, Paraguay

⁴Departamento de Física, CCNE, Universidade Federal de Santa Maria, Santa Maria, 97105-900, Brazil

⁵DCET, Universidade Estadual de Santa Cruz, Ilheus, 45662-900, Brazil

SDSS Data Release 7 (Abazajian et al. 2009)

0.01 < z < 0.3
mag > 17.78
54,896 objects
SDSS-DR7 / GZ1

<https://doi.org/10.1093/mnras/stad3934>

Monthly Notices
of the
ROYAL ASTRONOMICAL SOCIETY
MNRAS **528**, 82–107 (2024)
Advance Access publication 2023 December 21

Unveiling galaxy morphology through an unsupervised-su approach

I. Kolesnikov,^{1,2} V. M. Sampaio,^{3,4} R. R. de Carvalho,⁴ C. Conselice,² S. B. C. L. Mendes¹ and R. R. Rosa¹

¹Instituto Nacional de Pesquisas Espaciais, São José dos Campos, 12227-010 SP, Brazil

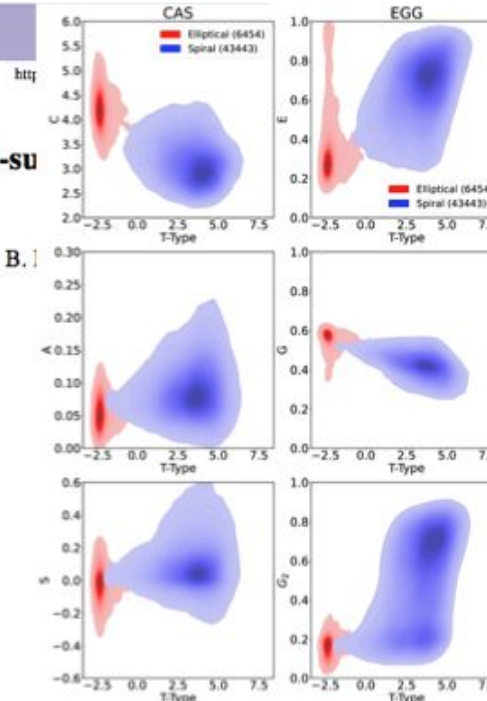
²Jodrell Bank Centre for Astrophysics, University of Manchester, Oxford Road, Manchester M13 9PL, UK

³School of Physics and Astronomy, University of Nottingham, University Park, Nottingham NG7 2RD, UK

⁴NAT - Universidade Cidade de São Paulo, 01506-000 SP, Brazil

⁵Universidade Federal de Santa Maria, Santa Maria, RS 97105-900, Brazil

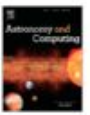
Dominguez Sánchez et al. (2018)
670,602 objects (+ T-Type Value)



<https://doi.org/10.1016/j.ascom.2019.100334>



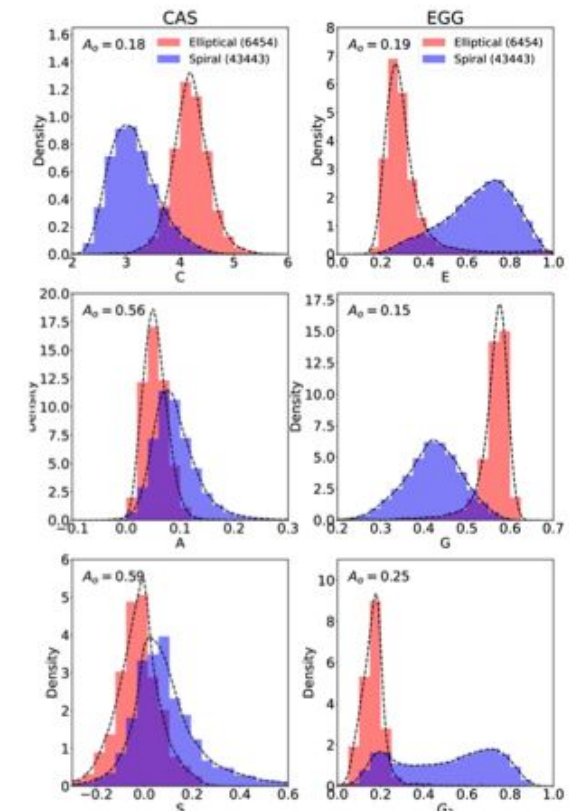
Astronomy and Computing
Volume 30, January 2020, 100334



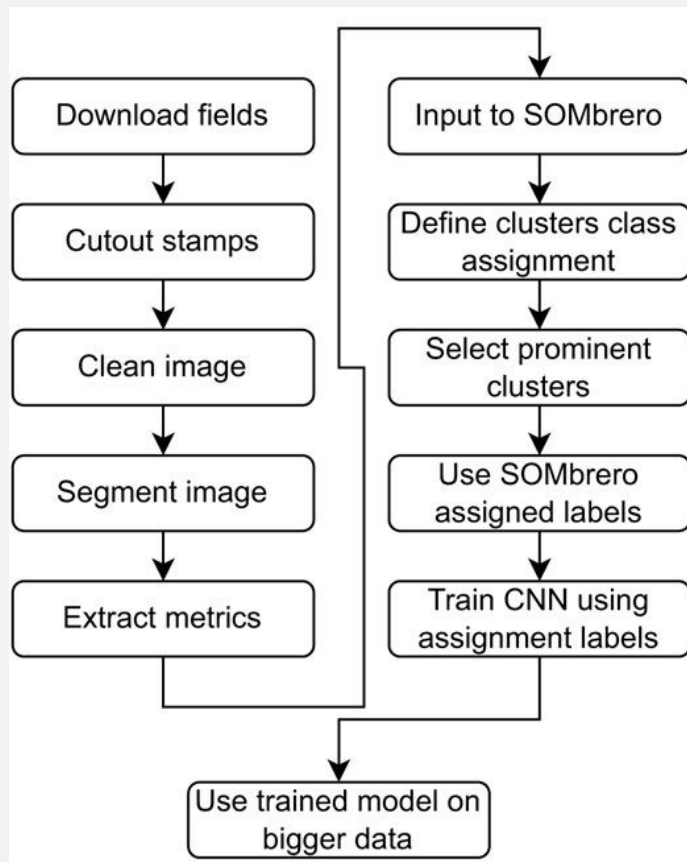
Full length article

Machine and Deep Learning applied to galaxy morphology - A comparative study

P.H. Barchi^{a, b}, R.R. de Carvalho^{c, d}, R.R. Rosa^a, R.A. Sautter^a,
M. Soares-Santos^b, B.A.D. Marques^e, E. Clua^e, T.S. Gonçalves^f,
C. de Sá-Freitas^f, T.C. Moura^g



Cymorph 3.0



Unveiling galaxy morphology through an unsupervised-supervised hybrid approach

I. Kolesnikov^{1,2*}, V. M. Sampaio^{3,4}, R. R. de Carvalho⁴, C. Conselice², S. B. Rembold⁵,
C. L. Mendes¹ and R. R. Rosa¹

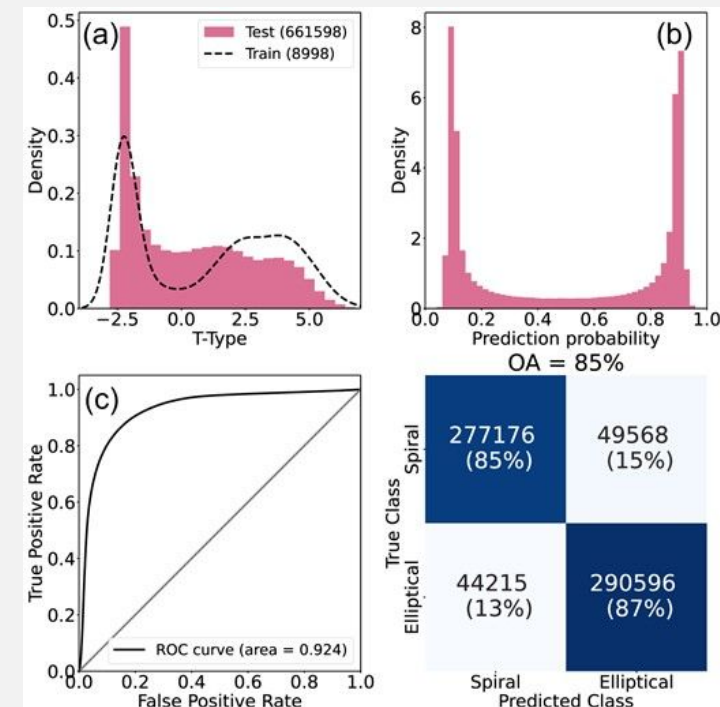
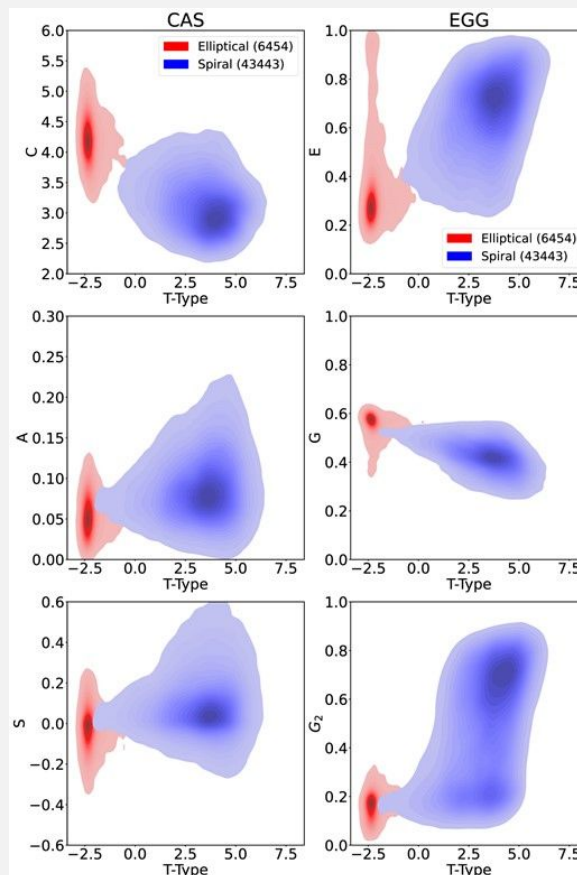
¹Instituto Nacional de Pesquisas Espaciais, São José dos Campos, 12227-010 SP, Brazil

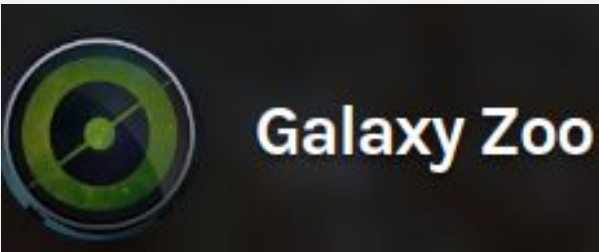
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⁴NAT - Universidade Cidade de São Paulo, 01506-000 SP, Brazil

⁵Universidade Federal de Santa Maria, Santa Maria, RS 97105-900, Brazil





How can generative AI positively impact the morphological classification of Galaxies?

The role of generative AI in training a *hyper citizen bot* for a Galaxy Zoo-like environment

https://github.com/indiradutta/DC_GAN

Deep Convolutional Generative Adversarial Network



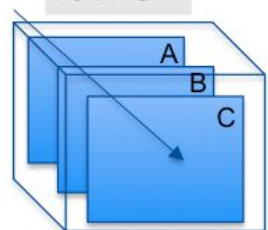
junyanz/**pytorch-CycleGAN-and-pix2pix**

Image-to-Image Translation in PyTorch

SDSS-Generator

Citizen-bot-for-Galaxies

Cb4Gx

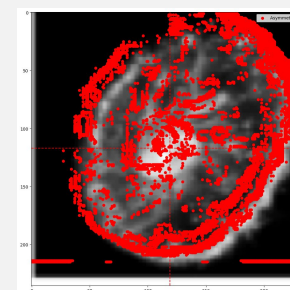
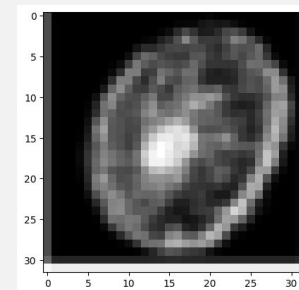
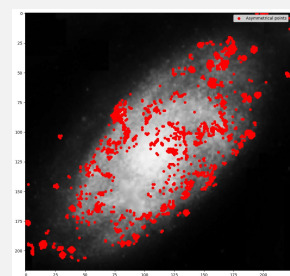
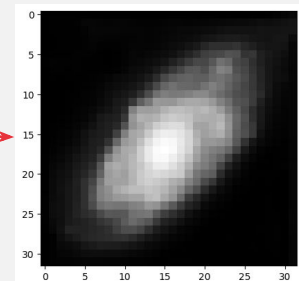
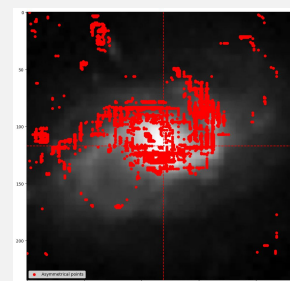
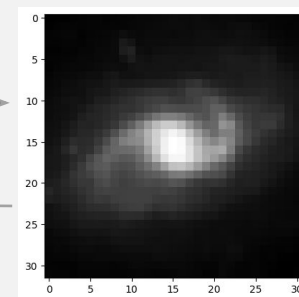
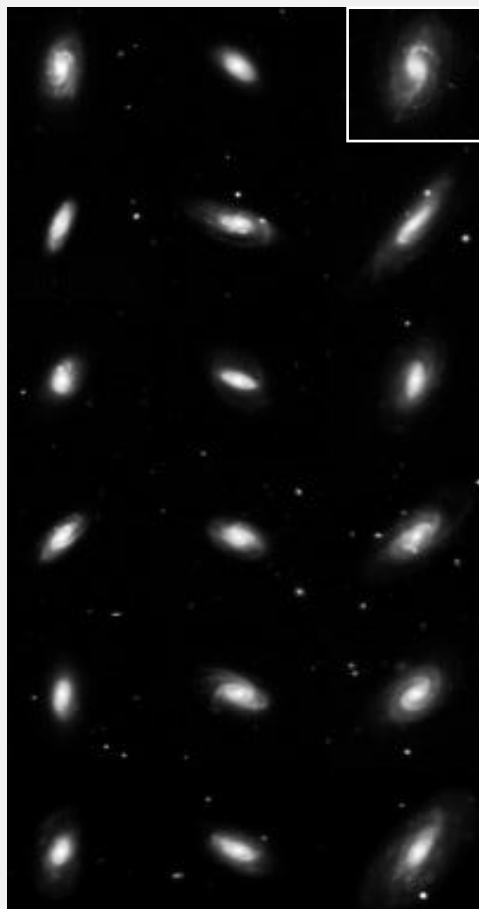


Polar Ring

Cb4Gx

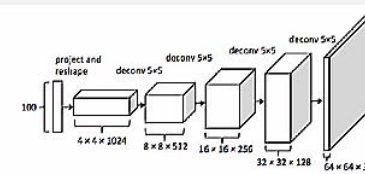


From-Any-Telescope

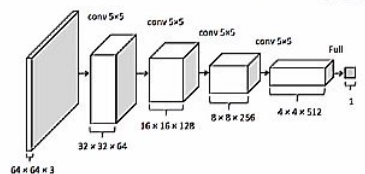


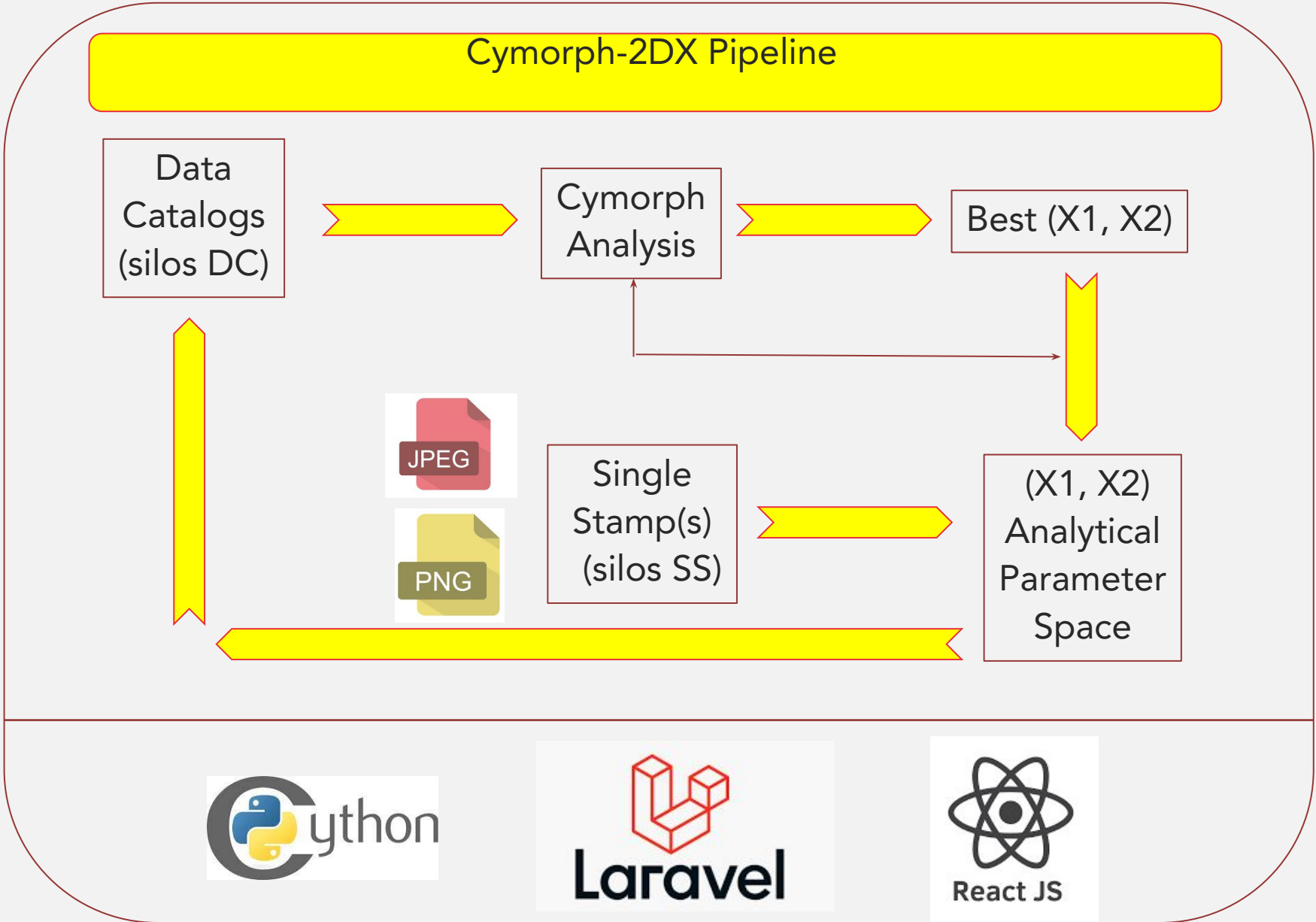
DCGAN Overall

Generator

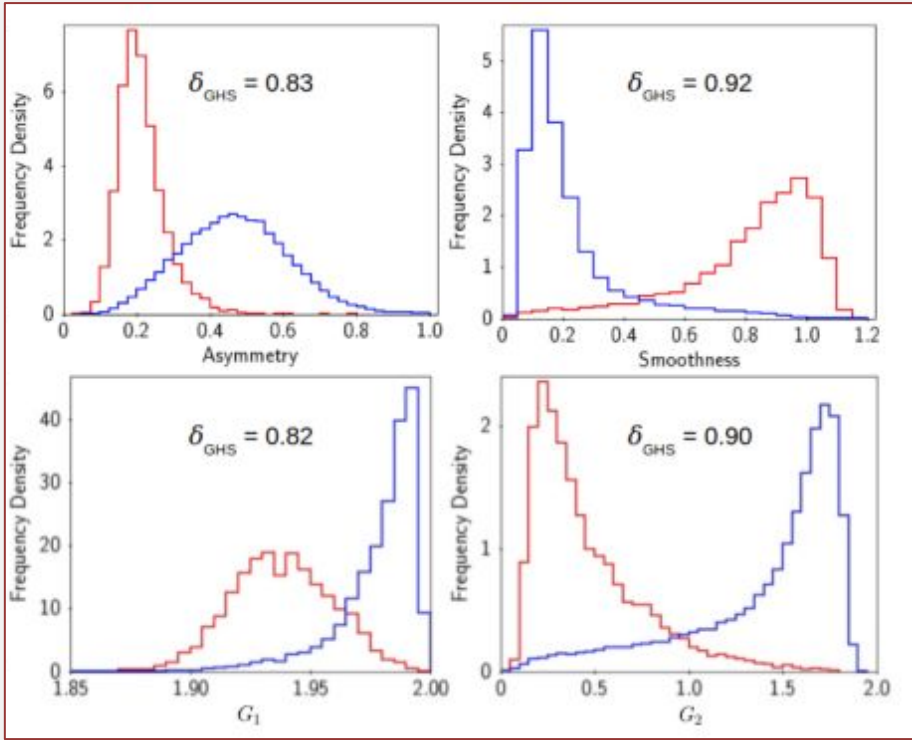
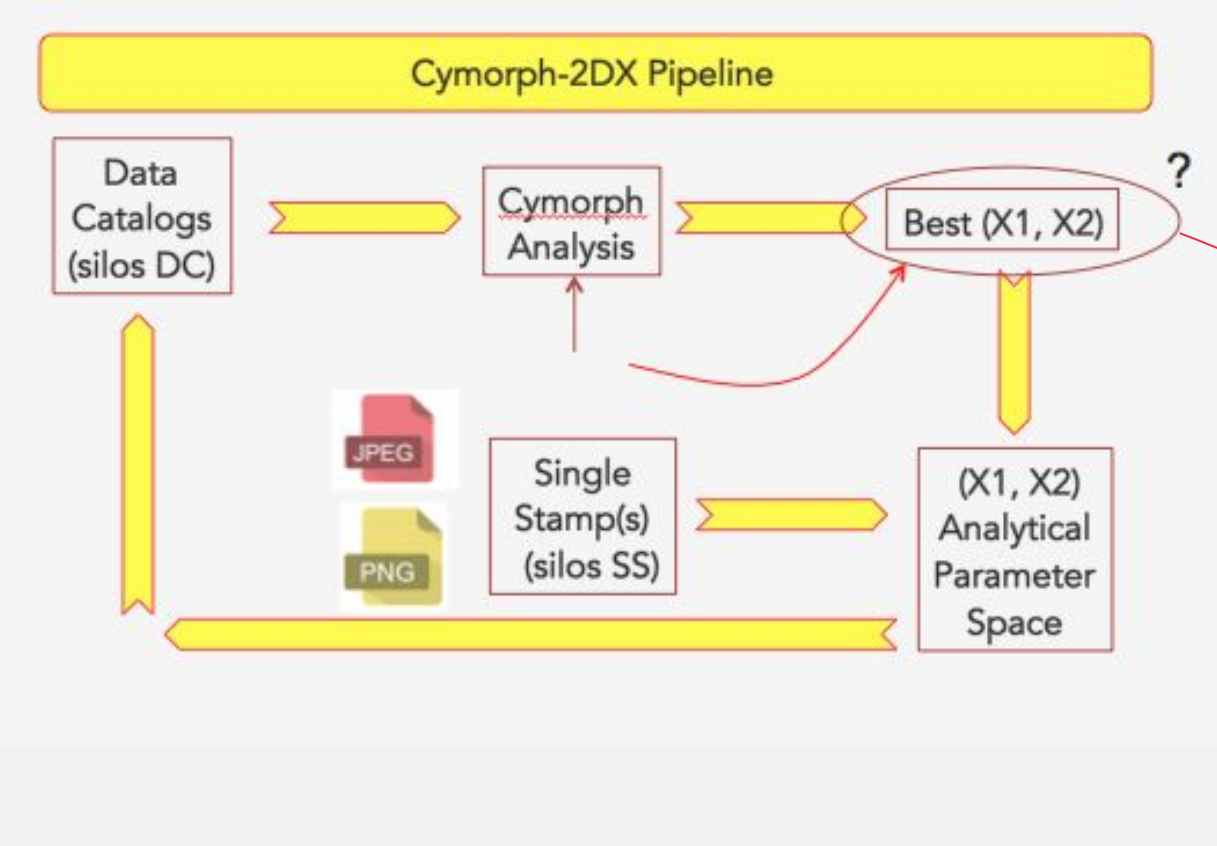


Discriminator





- Current Pipeline

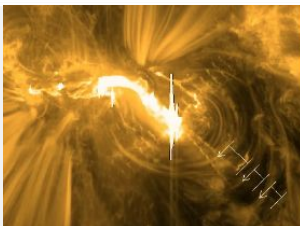
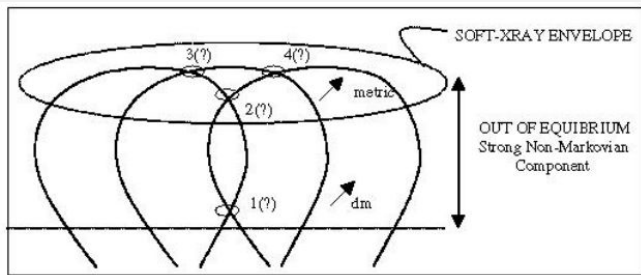
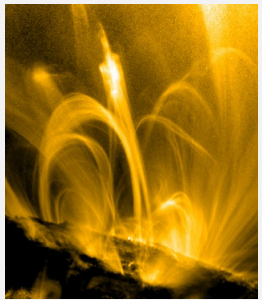


$$\delta_{\text{GHS}} = \frac{\left(1 - \frac{A_{\text{BR}}}{A_{\text{B}} + A_{\text{R}} + A_{\text{BR}}}\right)^{1/2} + \left(\frac{h_{\text{B}} + h_{\text{R}} - 2h_{\text{BR}}}{h_{\text{a}} + h_{\text{b}}}\right)}{2}$$

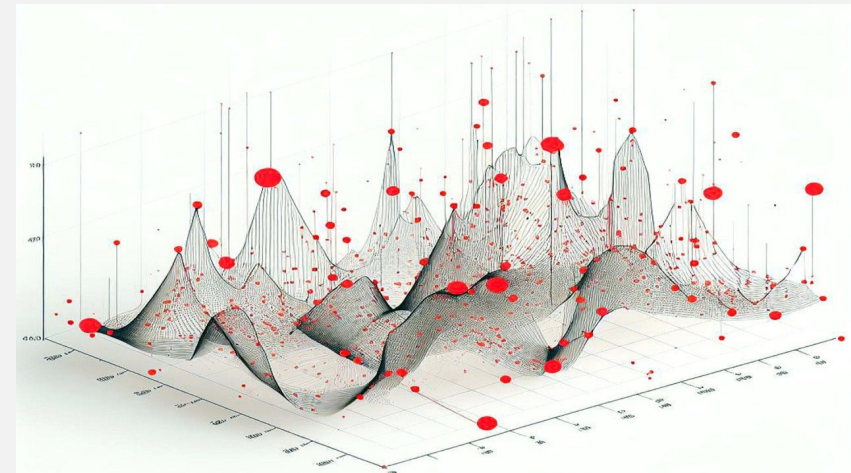
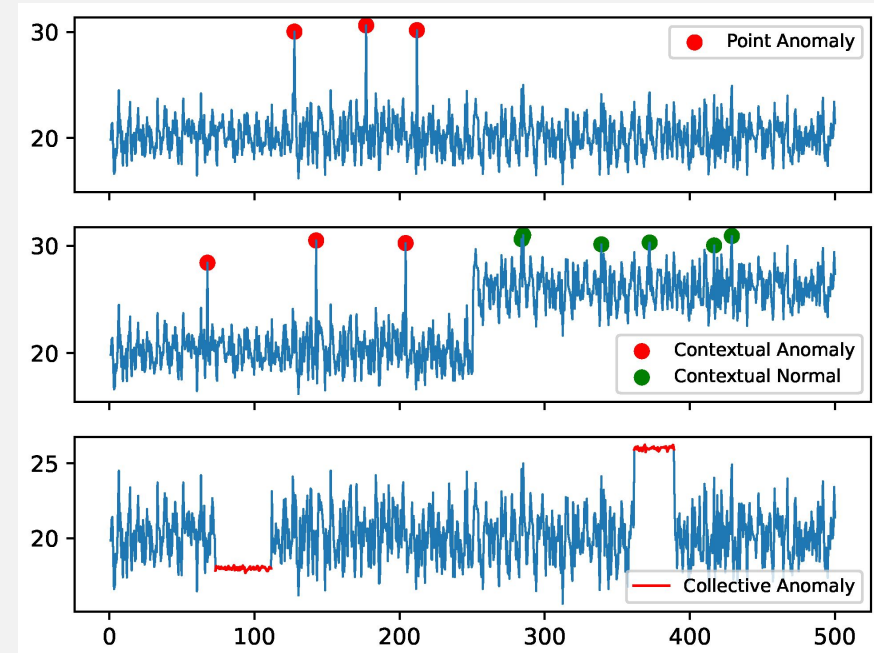
Rosa et al., MNRAS 477, L101-L105(2018)
doi:10.1093/mnras/sly054

Novo Módulo (ACDeep): Detecção e previsão de Anomalias

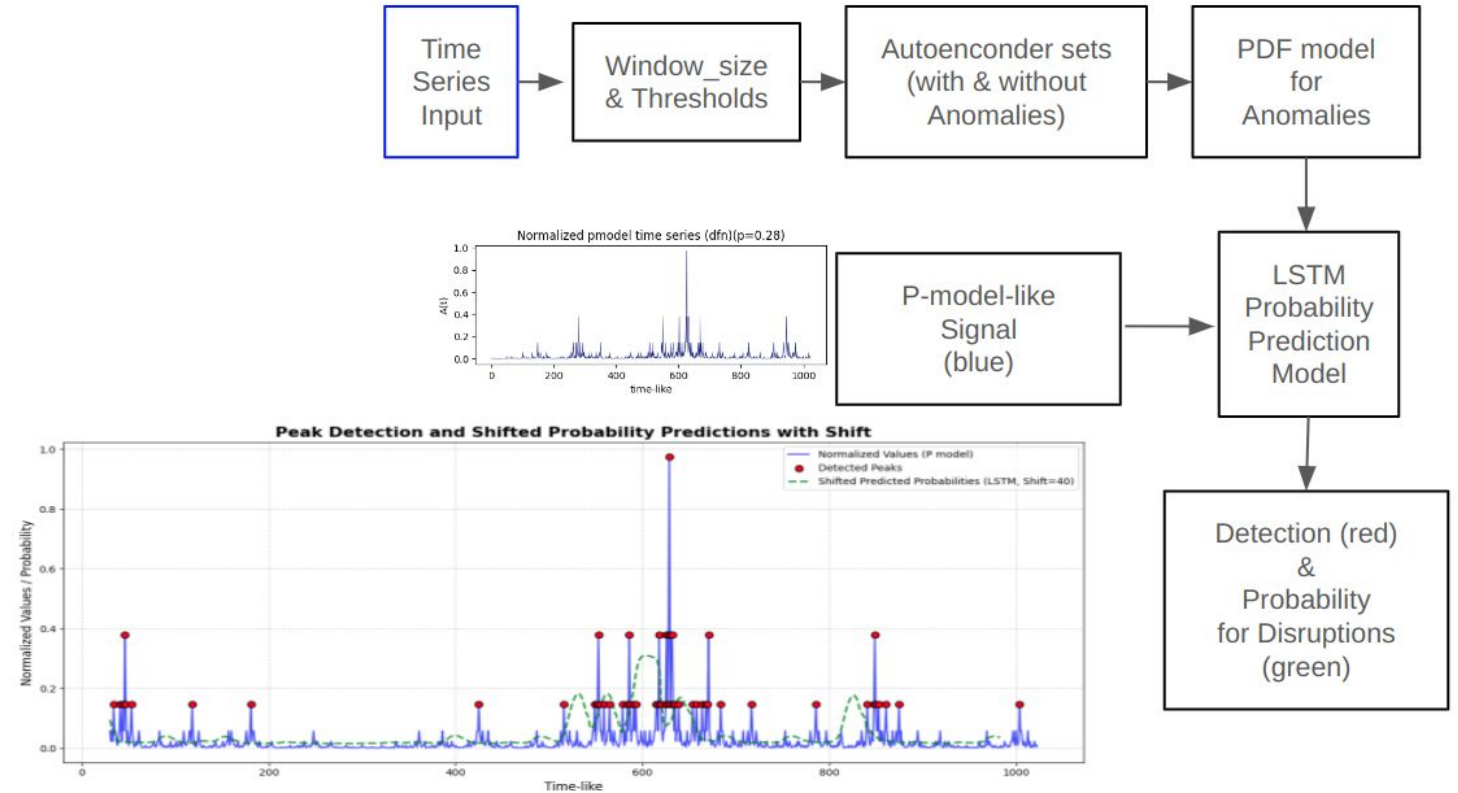
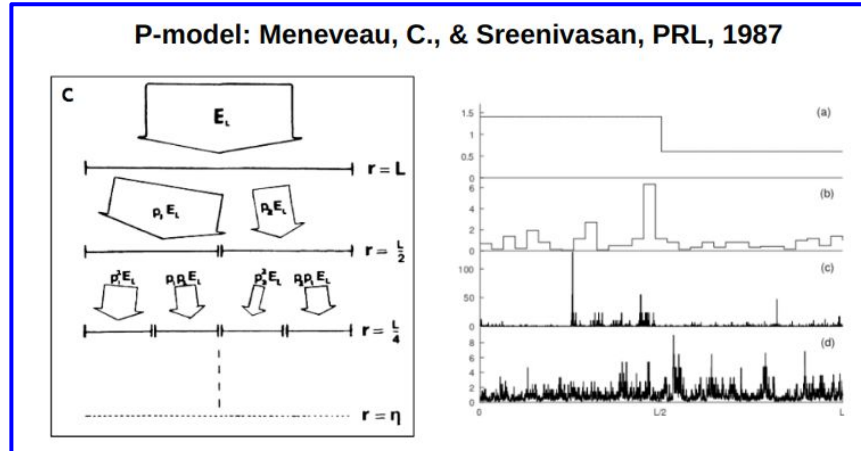
1. Disruption in Plasmas
2. Machine Learning Approaches
3. Data & ML Analysis
4. Results
5. Concluding Remarks



Disruption/Extremes = “Anomalies”



A LSTM-AUTOENCODER was trained and tested with data generated by p-model.

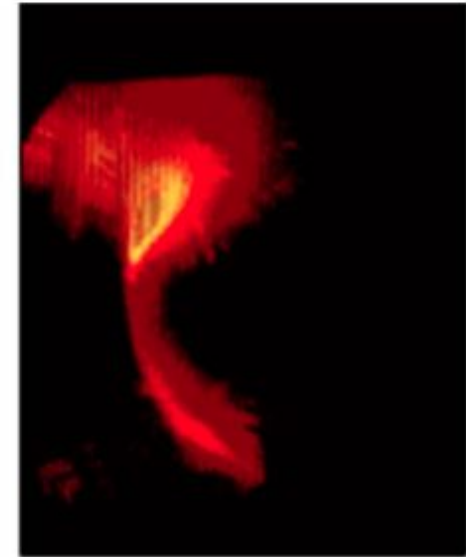
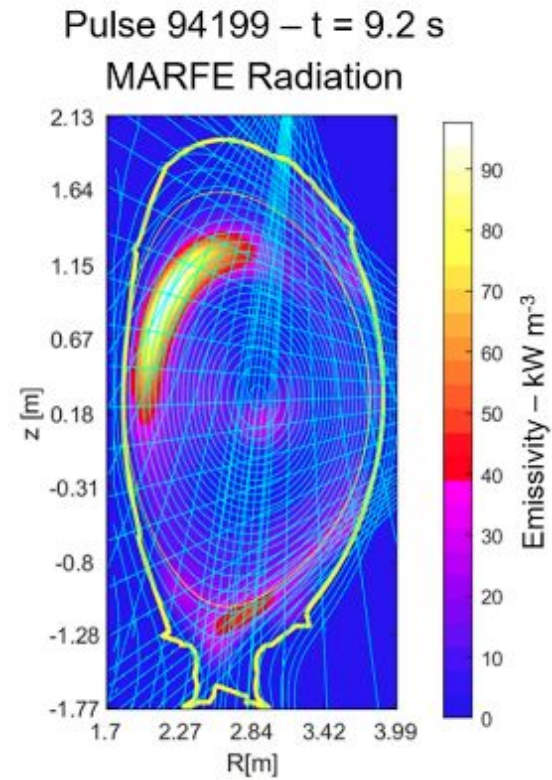
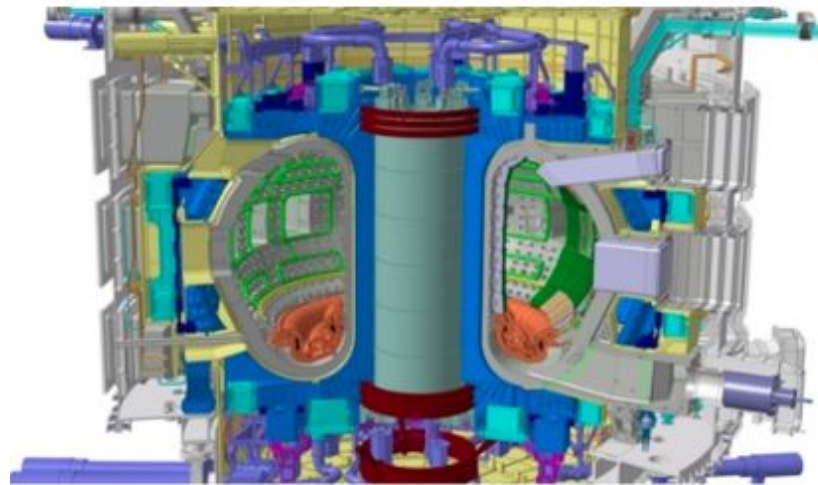


```
class EVTweibullAnomalyDetector:
    def __init__(self, window_size, threshold_factor, anomaly_threshold):
        self.window_size = window_size
        self.threshold_factor = threshold_factor
        self.anomaly_threshold = anomaly_threshold
        self.extreme_values = []
        self.weibull_params = None
```

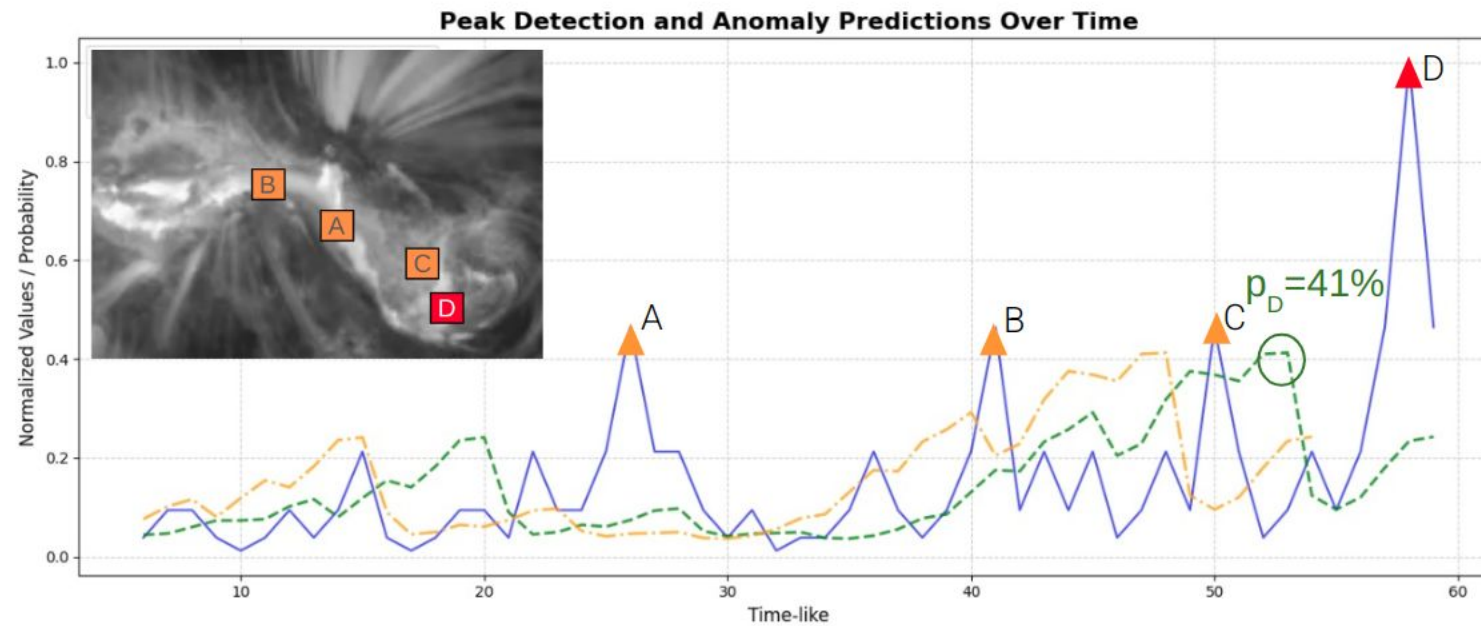
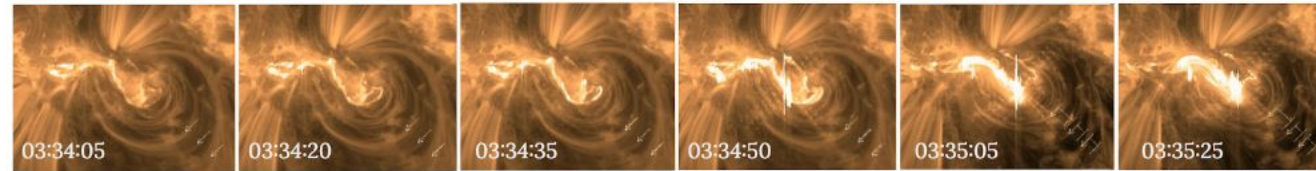
A control oriented strategy of disruption prediction to avoid the configuration collapse of tokamak reactors

[Andrea Murari](#), [Riccardo Rossi](#), [Teddy Craciunescu](#), [Jesús Vega](#), [JET Contributors](#) & [Michela Gelfusa](#) ✉

Nature Communications **15**, Article number: 2424 (2024) | [Cite this article](#)



Snapshots from SDO/171A - NOAA 11429 (09/3/12)
03:35:05 to 03:35:25 UT (180 frames)



Solar Physics & Space Weather

Automatically diagnose minor anomalies across multiple channels (or pixels) as quickly as possible to predict a larger systemic anomaly (disruption: a solar flare $\Rightarrow$ solar noise storms $\Rightarrow$ magnetic storm).

```
[3] #SPECIAL TOOLS FOR LSTM-AUTOENCODER - FOR - ANOMALIES PROBABILITIES  
#C. FALANDES-V.10-SEP-2024
```

```
from scipy.stats import weibull_min  
from tensorflow.keras.models import Model  
from tensorflow.keras.layers import Input, Dense  
import numpy as np  
import tensorflow as tf  
from tensorflow.keras.models import Sequential  
from tensorflow.keras.layers import LSTM, Dense, RepeatVector, TimeDistributed
```

```
[4] # Define autoencoder  
input_dim = 1024 # Example input size  
input_layer = Input(shape=(input_dim,))  
encoded = Dense(128, activation='relu')(input_layer)  
decoded = Dense(input_dim, activation='sigmoid')(encoded)  
autoencoder = Model(inputs=input_layer, outputs=decoded)  
autoencoder.compile(optimizer='adam', loss='mse')
```

```
[ ] class EVTWeibullAnomalyDetector:  
    def __init__(self, window_size, threshold_factor, anomaly_threshold):  
        self.window_size = window_size  
        self.threshold_factor = threshold_factor  
        self.anomaly_threshold = anomaly_threshold  
        self.extreme_values = []  
        self.weibull_params = None
```

CyMorph + CyMorph2DX + ACDeep



Rubens



Luan



Carlos E. Falandes



Welcome to
Jupyter Lab



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Outline:

- Galaxies (LSST era)
- Morphometrics (Fundamentals)
- Morphometrics (Strategies)**
- Concluding Remarks

